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# Optical encoder based on the fractional Talbot effect

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#### Abstract

This paper presents a reflective optical encoder based on the Talbot self-image of binary phase grating. The proposed encoder has several advantages in comparison to conventional Moire´ effect based encoders, including larger gap between gratings, higher gap tolerance and greater light passage efficiency. This encoder with gap 220  $\mu$ m can accurately measure the position of linear stages which move at high-speed. The encoder includes a grating scale with period 20  $\mu$ m and provides a signal period of 20  $\mu$ m with a frequency response of approximately 200 kHz. This system and the associated electronics provide an interpolation factor of 200 that gives a resolution of 0.1  $\mu$ m. The repeatability is 0.2  $\mu$ m for a traveling length of 600 mm.

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# 1. Introduction

Most optical encoders (linear transducer as well as rotational encoder) are based on the Moiré effect. A collimated beam penetrates two gratings, which have equal period. Then, an oscillating signal is generated when a grating moves in the direction which is perpendicular to its ruling. In actual optical encoders, the most commonly used light source is an infrared LED with a typical size of 350  $\mu$ m. The optics used to collimate the light beam typically comprises a high-numerical-aperture singlet with a focal length of about 10 mm, due to the limitations on the total size of the reading head of the encoder. With this ratio of source size to focal length, the divergent angle of light is around  $2^{\circ}$ . The contrast decay of the signal can be described

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<span id="page-1-0"></span>by  $m = (2g/T) \tan \gamma$ , where g denotes the gap between two gratings,  $\gamma$  is the divergent angle of light, and T is the grating period [\[1\]](#page-7-0). For example, for an encoder mentioned above and with a grating period of 20  $\mu$ m, the gap between gratings should be roughly 50  $\mu$ m in order to get a good signal. Such a small gap between gratings could cause a big problem in mechanical tolerance in the manufacture of these encoders. In the Moiré type encoders, a bearing and spring are used to maintain the gap. However, the bearing cannot work for high-speed motion. Meanwhile, in high-speed motion, a reflective type of encoder can be used without bearing. Crespo et al. [\[2\]](#page-7-0) designed an optical encoder using the Lau effect in a reflective-type encoder to increase the gap. Nevertheless, they used an amplitude-type grating, through which only half of light flux can pass through.

This paper presents an optical encoder using the Talbot effect [\[3–5\].](#page-7-0) A phase grating which produces a self-image at a quarter Talbot distance is used in the encoder [\[6–8\]](#page-7-0), so that all light can penetrate it. This reflective-type encoder with phase grating has the advantages of a large gap and more light efficiency. The remainder of this paper is organized as follows: In Section 2.1, we derive expressions for the self-image of phase grating. In Section 2.2, we discuss a reflective-type optical encoder, using the three-grating system model. In Section 3, we present some measurements made using the reflective-type encoders and evaluate the manufacturing tolerance.

## 2. Theory

## 2.1. Fractional Talbot effects of gratings

Self-imaging properties were first observed by Talbot in 1836. During the 1960s, Winthrop and Worthington investigated many theoretical and practical aspects, especially for fractional Talbot planes. In this study, we applied the theory of Winthrop and Worthington in a one-dimensional case. The transmittance of a one-dimensional periodic structure shown in Fig. 1 with unit cell transmittance  $t_p(x)$  can be described by

$$
t(x) = t_p(x) \otimes \text{comb}\left(\frac{x}{p}\right),\tag{1}
$$

where p denotes the spatial period in the x-direction and  $\otimes$  represents convolution. From Fresnel diffrac-tion [\[9\]](#page-7-0), the optical field in an arbitrary plane at a distance  $z_F$  behind the periodic structure is given by

$$
U(x, z_{F}) = C(z_{F}) \times t_{p}(x) \otimes \left\{ \sum_{m=-\infty}^{\infty} \exp\left[-i\pi \lambda z_{F}\left(\frac{m}{p}\right)^{2}\right] \exp\left[i2\pi \left(\frac{m}{p}\right)x\right]\right\},
$$
\n(2)

Fig. 1. One-dimension binary grating with period P.

<span id="page-2-0"></span>where  $C(z_F) = (\exp(i k z_F) / i \lambda z_F) \cdot \sqrt{2z_F/k} \exp(i\pi/4)$  and  $\lambda$  is the light wavelength. In this study, the proposed encoder is based on the fractional Talbot effect and only the diffraction amplitude on the plane at  $z = z_T/4$  is considered, where  $z_T$  denotes the Talbot distance and equals  $2p^2\lambda$ . Accordingly, Eq. [\(2\)](#page-1-0) can be rewritten as

$$
U\left(x,\frac{z_T}{4}\right) = A \times \left\{ t_p(x) \otimes \sum_{m=-\infty}^{\infty} \exp\left(-i\pi \frac{m^2}{2}\right) \exp\left[i2\pi \left(\frac{m}{p}\right)x\right] \right\},\tag{3}
$$

where  $A = \sqrt{2}/(ip\sqrt{\pi}) \exp[i\pi (p/\lambda)^2 + i\pi/4]$ . By setting  $m = 2n_0 + n_1$ , the summation term on the right-hand side of Eq. (3) can be rearranged as

$$
\sum_{m=-\infty}^{\infty} \exp\left(-i\pi \frac{m^2}{2}\right) \exp\left[i2\pi \left(\frac{m}{p}\right)x\right] = \sum_{n_0=-\infty}^{\infty} \sum_{n_1=0,1} \exp\left(-i2\pi n_0^2\right) \exp\left(-i2\pi n_0 n_1\right) \times \exp\left[-i\pi \left(\frac{n_1^2}{2}\right)\right] \exp\left[i4\pi \left(\frac{n_0}{p}\right)x\right] \exp\left[i2\pi \left(\frac{n_1}{p}\right)x\right].
$$
\n(4)

The terms  $\exp(-i2\pi n_0^2)$  and  $\exp(-i2\pi n_0n_1)$  in Eq. (4) equals one because  $n_0$  and  $n_1$  are integers. Thus, Eq. (4) can be further simplified as

$$
\sum_{m=-\infty}^{\infty} \exp\left(-i\frac{\pi}{2}m^2\right) \exp\left[i2\pi\left(\frac{m}{p}\right)x\right] = \left\{1 + \exp\left[i\frac{2\pi}{p}\left(x - \frac{p}{4}\right)\right]\right\} \times \sum_{n_0=-\infty}^{\infty} \delta\left(x - n_0\frac{p}{2}\right)
$$

$$
= \sum_{n_0=-\infty}^{\infty} \left\{1 + \exp\left[i\pi\frac{(2n_0 - 1)}{2}\right]\right\} \delta\left(x - n_0\frac{p}{2}\right),\tag{5}
$$

where  $\delta$  represents the Dirac delta function. Substituting Eq. (5) into Eq. (3), the amplitude distribution  $U(x, z_T/4)$  becomes

$$
U\left(x,\frac{z_{\rm T}}{4}\right) = A \sum_{n_0=-\infty}^{\infty} \left\{ \left[1 + \exp\left(i\pi \frac{(2n_0-1)}{2}\right)\right] \times t_p\left(x-\frac{n_0p}{2}\right) \right\}.
$$
 (6)

The parity of  $n_0$  determines the value of the phase term in Eq. (6). If  $n_0$  is odd, then  $\{1 + \exp[i\pi(2n_0 - 1)]$ 2]} = (1 + i) and  $n_0 = 2n + 1$ . On the contrary, if  $n_0$  is even, then  $\{1 + \exp[i\pi(2n_0 - 1)/2]\} = (1 - i)$  and  $n_0 = 2n$ . These results lead Eq. (6) to

$$
U\left(x, \frac{z_1}{4}\right) = A \sum_{n=-\infty}^{\infty} \left[ (1-i)t_p(x - np) + (1+i)t_p\left(x - np - \frac{p}{2}\right) \right]
$$
  
=  $\sqrt{2}A \exp\left(-i\frac{\pi}{4}\right) \left[ t(x) + \exp\left(i\frac{\pi}{2}\right)t\left(x - \frac{p}{2}\right) \right].$  (7)

Eq. (7) indicates that the field distribution in the quarter Talbot plane [\[7–9\]](#page-7-0) consists of two parts. The both parts are the same as the transmittance of the periodic object. But there exist a displacement of a half period and a phase shift of  $\pi/2$  between the two parts.

In this study, the proposed encoder contains two different periodic structures: binary-phase grating (BPG) and binary-amplitude grating (BAG). Supposing that these gratings have the same periodicity, the phase step of the BPG is  $\phi$ , and the open ratio of the BAG is 0.5. The transmittances of the BPG and BAG can be expressed as follows:

$$
t_{\rm BPG}(x) = \left\{ \exp\left(i\phi\right) \operatorname{rect}\left(\frac{x - \frac{p}{4}}{\frac{p}{2}}\right) + \operatorname{rect}\left(\frac{x - \frac{3}{4}p}{\frac{p}{2}}\right) \right\} \otimes \operatorname{comb}\left(\frac{x}{p}\right),\tag{8}
$$

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<span id="page-3-0"></span>
$$
t_{\text{BAG}}(x) = \text{rect}\left(\frac{x - \frac{p}{4}}{\frac{p}{2}}\right) \otimes \text{comb}\left(\frac{x}{p}\right). \tag{9}
$$

If the input object is a binary-phase grating, i.e.,  $t(x) = t_{BPG}(x)$  and the phase step  $\phi = \pi/2$ , then the field distribution on the  $z_T/4$  plane, obtained by substituting Eq. [\(8\)](#page-2-0) into Eq. [\(7\),](#page-2-0) is

$$
U_{\rm BPG}\left(x,\frac{z_{\rm T}}{4}\right) = 2\sqrt{2}A\exp\left(i\frac{\pi}{4}\right)\operatorname{rect}\left(\frac{x-\frac{p}{4}}{\frac{p}{2}}\right)\otimes\operatorname{comb}\left(\frac{x}{p}\right). \tag{10}
$$

According to Eq. (9), the above expression shows a binary-amplitude grating located at a quarter Talbot distance. This lens-like imaging effect can be reversed if the input object is a binary-amplitude grating, i.e.,  $t(x) = t_{BAG}(x)$ . By substituting Eq. (9) into Eq. [\(7\),](#page-2-0) the field distribution on the quarter Talbot plane is

$$
U_{\text{BAG}}\left(x,\frac{z_{\text{T}}}{4}\right) = \sqrt{2}A\exp\left(-i\frac{\pi}{4}\right)\left[\text{rect}\left(\frac{x_0-\frac{p}{4}}{\frac{p}{2}}\right) + \exp\left(-i\frac{\pi}{2}\right)\text{rect}\left(\frac{x_0-\frac{3p}{4}}{\frac{p}{2}}\right)\right] \otimes \text{comb}\left(\frac{x}{p}\right). \tag{11}
$$

#### 2.2. Three-grating model  $[10]$

Fig. 2 depicts the structure of the optical linear encoder presented in this study. The binary-phase and binary-amplitude gratings are the index and main scales, respectively. Both gratings have the same period and their separation is a quarter of the Talbot distance. The collimated light passes through the index scale and produces a BAG image on the main scale, as described in Section 2.1. The light then is reflected by the main scale and goes through the index scale again. Finally, a photo-detector collects the reflected light. The output current of the detector changes with the displacement between the index grating and the main scale.

In the encoder, the light received by the detector passes through the index grating twice and is reflected by the main grating once. [Fig. 3 illustrates that the optical configuration is equivalent to a three-grating](#page-4-0) [system,](#page-4-0) where  $G_1$  and  $G_3$  are the index grating and  $G_2$  is the main grating. A monochromatic plane wave is diffracted by  $G_1$  and propagates to  $G_2$ . Because the distance between  $G_1$  and  $G_2$  is  $z_T/4$ , the diffraction filed in front of  $G_2$  is given by Eq. (10),

$$
U_{2f}(x) = 2\sqrt{2}A \exp\left(-i\frac{\pi}{4}\right) \operatorname{rect}\left(\frac{x-\frac{p}{4}}{\frac{p}{2}}\right) \otimes \operatorname{comb}\left(\frac{x}{p}\right). \tag{12}
$$

The above equation shows that the field diffracted by  $G_1$  is the same as the transmittance of grating  $G_2$ . The self-image of  $G_1$  overlaps on grating  $G_2$ , therefore, the Moiré fringe is generated when one grating moves in



Fig. 2. Schematic diagram for the reflective optical encoder.

<span id="page-4-0"></span>

Fig. 3. The three-grating system.

the x-direction, as shown in Fig. 3. If a displacement  $\Delta x$  occurs between  $G_1$  and  $G_2$ , the diffraction field just behind  $G_2$  is

$$
U_{2b}(x) = U_{2f}(x) \times t_{BAG}(x) = \text{rect}\left(\frac{x - \frac{p}{4} - \Delta x}{\frac{p}{2} - \Delta x}\right) \otimes \text{comb}\left(\frac{x}{p}\right).
$$
\n(13)

The Moiré fringe given by Eq.  $(13)$  seems like the field that a plane wave just passes through a binary-amplitude grating laterally moved by  $\Delta x$ . Additionally, the open ratio of the grating is changed to  $(0.5 - \Delta x/p)$ . The wave  $U_{2b}(x)$  keeps propagating and arrives at G<sub>3</sub>. The description of the wave  $U_{3f}(x)$  before G<sub>3</sub> is sim-ilar to Eq. [\(11\)](#page-3-0) because the travel distance equals  $z_T/4$ . Hence,  $U_{3f}(x)$  is given by

$$
U_{3f}(x) = A_2 \exp\left(-i\frac{\pi}{4}\right) \left[ \operatorname{rect}\left(\frac{x - \frac{p}{4} - \frac{\Delta x}{2}}{\frac{p}{2} - \Delta x}\right) \exp\left(-i\frac{\pi}{2}\right) + \operatorname{rect}\left(\frac{x - \frac{3p}{4} - \frac{\Delta x}{2}}{\frac{p}{2} - \Delta x}\right) \right] \otimes \operatorname{comb}\left(\frac{x}{p}\right). \tag{14}
$$

After passing  $G_3$ , the field distribution becomes

$$
U_{3b}(x) = U_{3f}(x) \times t_{\text{BPG}}(x) = A_2 \exp\left(-i\frac{\pi}{4}\right) \left[ \text{rect}\left(\frac{x - \frac{p}{4} - \frac{\Delta x}{2}}{\frac{p}{2} - \Delta x}\right) + \text{rect}\left(\frac{x - \frac{3p}{4} - \frac{\Delta x}{2}}{\frac{p}{2} - \Delta x}\right) \right] \otimes \text{comb}\left(\frac{x}{p}\right). \tag{15}
$$

Finally, the light wave reaches the detector. Since the size of the detector is much larger than the grating period, the detector output can be expressed by

$$
S(\Delta x) \propto \int_0^p |U_{3b}(x)|^2 dx \propto A\left(\frac{\Delta x}{p}\right),\tag{16}
$$

where  $A$ (argument) is the triangle function. Eq. (16) shows that the detector output is a triangle function of  $\Delta x$ , which is the lateral displacement between  $G_1$  and  $G_2$ . However, in practice the shape of the output signal is similar to a sinusoidal function owing to the non-perfect collimated light and finite grating structure.

## 3. Experimental setup

[Fig. 2](#page-3-0) schematically depicts a setup for the reflective-optical encoder. In this experiment, an infrared light emitted diode (LED) is used as light source. Though laser diode (LD) has better high properties than LED, we consider the compactness of the head and the LED meets all the requirements for signal to noise ratio (SNR) and contrast in this case.

The LED light source with a wavelength of  $0.88 \mu m$  is collimated by a collimating lens with focal length of 9 mm. The collimated beam passes through the phase grating and is reflected by grating  $G_2$ . The phasegrating  $G_1$  and reflective-grating  $G_2$  have the same period 20  $\mu$ m. The phase grating is divided into four segments. The aperture of segment is 1 mm; consequently, the pupil size covers the period number of 50. The number of periods is sufficient to product the self-imaged effect. The spacing between adjacent segments is  $(20N + 5)$  µm, where N is the integer, as shown in Fig. 4. The four segments in the phase grating  $G_1$ correspond to four sinusoidal outputs with phase shifts by  $90^{\circ}$  from four photo-detectors. The phase grating  $G_1$  is illuminated with the collimated light and self-imaged on grating  $G_2$ . A quarter of Talbot distance is 0.22 mm. The luminous dimension of LED is about  $160 \mu m$ . With this ratio of luminous dimension to focal length, the divergent angle of light source is about  $1^\circ$ . Fig. 5 shows the contrast variation of the Moiré signal due to the gap change. The normalized contrast  $m$  can be defined by

$$
m = \frac{\text{AC}(g)}{\text{AC}(0.22)},\tag{17}
$$

where  $AC(g)$  denotes the amplitude of the encoder signal when the gap between two gratings is g and AC(0.22) represents the amplitude of the encoder signal when the gap is 0.22 mm. For getting a good signal (the minimum contrast must be greater than 0.6), the tolerance of gap is approximately  $\pm 70 \text{ }\mu\text{m}$ .

From Eq. [\(7\)](#page-2-0), the binary phase grating is converted into a binary-amplitude grating image at the quarter of the Talbot distance when the phase step is  $\phi$ . The intensity distribution of the image is [\[11\]](#page-7-0)



Fig. 4. Scheme of four-segmented phase gratings. The zone spacing between adjacent segments is  $(20N + 5)$  µm, where N is integer. There is a 90° phase shift between adjacent segments.



Fig. 5. The contrast variation of signal due to the gap change.



Fig. 6. Two orthogonal signals (sine and cosine) from an oscilloscope.

$$
I = \left| U\left(x, \frac{z_1}{4}\right) \right|^2 = \left| \sqrt{2}A \exp\left(-i\frac{\pi}{4}\right) \left[ t\left(x - \frac{p}{2}\right) \exp\left(i\frac{\pi}{2}\right) + t(x) \right] \right|^2
$$
  
 
$$
\propto \text{rect}\left(\frac{x - \frac{p}{4}}{\frac{p}{2}}\right) \cdot (1 + \sin \phi) + \text{rect}\left(\frac{x - \frac{3p}{4}}{\frac{p}{2}}\right) \cdot (1 - \sin \phi). \tag{18}
$$

The step height of phase grating is  $h = \lambda \phi/2\pi(n - 1)$ . If manufacturing error  $\Delta h$  occurs for the step height, the phase shifts with  $\Delta \phi$ . The phase shift can be written as

$$
\Delta \phi = \frac{2\pi (n-1)\Delta h}{\lambda}.\tag{19}
$$

The visibility of the grating pattern can be calculated by

$$
V = \frac{I_{\text{max}} - I_{\text{min}}}{I_{\text{max}} + I_{\text{min}}} = \sin\left(\phi + \Delta\phi\right)
$$
\n<sup>(20)</sup>

If the phase step is  $\phi = \pi/2$ , the step height of phase grating is  $h = \lambda \phi/2\pi (n - 1) = 430$  nm, and the visibility can be expressed by  $cos(\Delta \phi)$ . The step depth of phase grating measured by the atomic force microscopy approximately equals  $h = 480$  nm and the error of manufacture is 50 nm, which corresponds to a phase shift of 0.0575 $\pi$ . According to Eq. (20), the visibility of self-image, V, equals 98%. The contrast of the Moiré signal between phase grating  $G_1$  and reflective grating  $G_2$  decreases from 100% to 98%. For a good visibility ( $>95\%$ ), the step depth tolerance of the phase grating is  $\pm 87$  nm.

A linear positioning system with a scanning range of 60 cm was used to measure the accuracy of the optical encoder. The linear positioning system is a linear motor with position feedback control by laser interferometer. In this study, a laser interferometer is used as a standard for measuring the encoder error. Fig. 6 shows two orthogonal signals (sine and cosine) with a signal period of  $20 \mu m$ . The resolution could be im-proved to 0.1 µm using interpolation circuit with a factor of 200. [Fig. 7](#page-7-0) shows the three-path (back and forth) repeatability measurement. The vertical axis represents the measuring difference between the optical encoder and the laser interferometer. Meanwhile, the horizontal axis is the travel distance of the linear stage. In this work, a repeatability error of  $0.5 \mu m$  is obtained that includes the backlash of the linear stage. We can divide the total repeatability into two parts – forward and backward. Both of them have repeatability error of less than  $0.2 \mu m$ .

<span id="page-7-0"></span>

Fig. 7. Three-path (back and forth) repeatability measurement. The vertical axis was the measuring difference between optical encoder and laser interferometer. The horizontal axis was the travel distance of linear stage. We have obtained a repeatability of  $0.5 \mu m$ .

# 4. Conclusion

In this work, we develop a reflective optical encoder using Talbot image of the binary phase grating. In comparison with conventional Moiré encoder, the Talbot effect provides a larger gap between two gratings. The phase grating has more light to be passed and gives higher signal intensity. This investigation provides the theoretical expressions for the gap tolerance and manufacture tolerance of phase grating. Experimental results show a good agreement with theoretical results. For high moving speed, the optical encoders have sinusoidal output signals particularly with levels of 1 V at  $a - 3$  dB cutoff frequency of approximately 200 kHz. For sub-micro resolution measuring steps, an electronic interpolation with a factor of 200 leads to a measuring resolution of 0.1  $\mu$ m. This study obtains excellent repeatability of 0.5  $\mu$ m and large gap of 0.22 mm between two gratings with a pitch of 20  $\mu$ m for 600 mm traveling range.

Further efforts will focus on improving the resolution and accuracy by using a fine grating period, such as 10  $\mu$ m. The gap between two gratings with a period of 10  $\mu$ m is about 60  $\mu$ m. In order to have a large gap, the system should work in a higher-order Talbot zone. The laser diode source will be used for higher brightness, better wave-front quality and narrower spectral bandwidth.

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